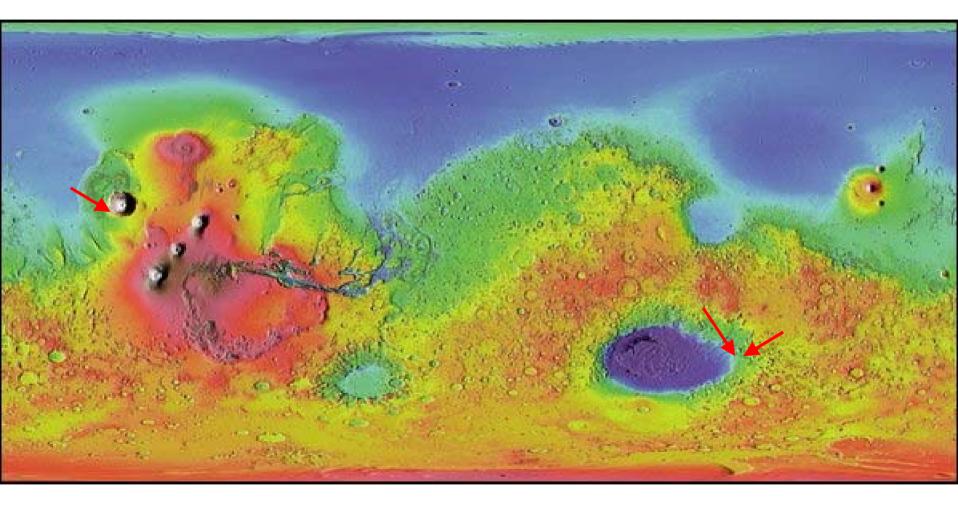
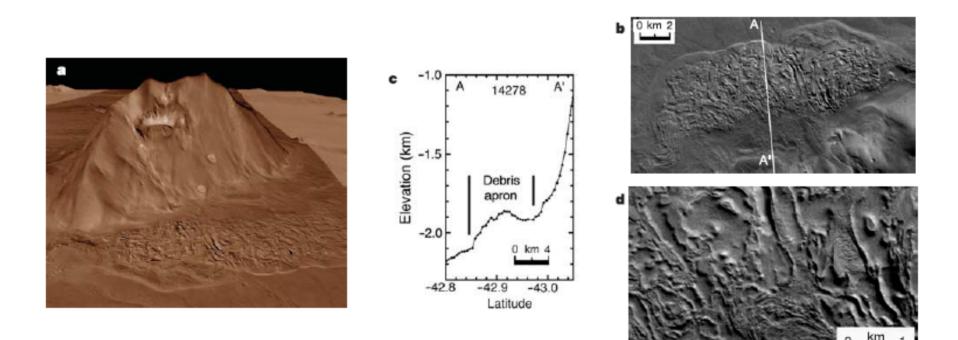
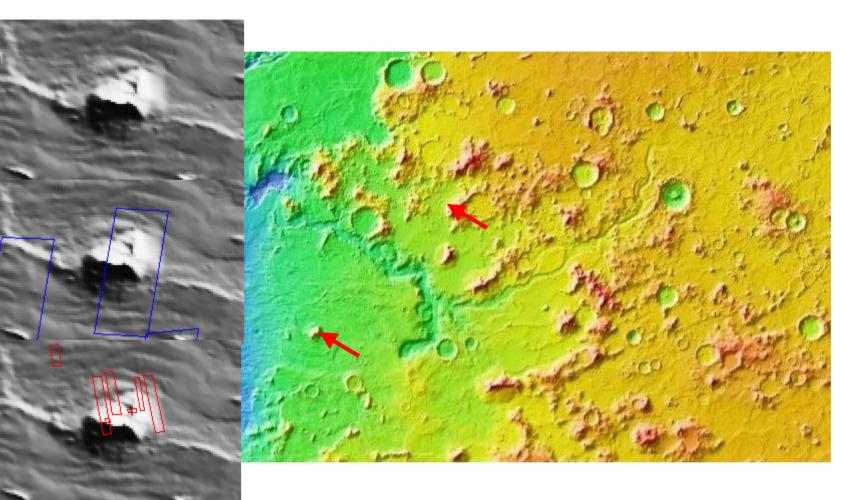
## **Tropical to mid-latitude snow and ice accumulation, flow and glaciation on Mars**

J. W. Head<sup>1</sup>, G. Neukum<sup>2</sup>, R. Jaumann<sup>3</sup>, H. Hiesinger<sup>1</sup>, E. Hauber<sup>3</sup>, M. Carr<sup>4</sup>, P. Masson<sup>5</sup>, B. Foing<sup>6</sup>, H. Hoffmann<sup>3</sup>, M. Kreslavsky<sup>1</sup>, S. Werner<sup>2</sup>, S. Milkovich<sup>1</sup>, S. van Gasselt<sup>2</sup> & The HRSC Co-Investigator Team<sup>\*</sup>

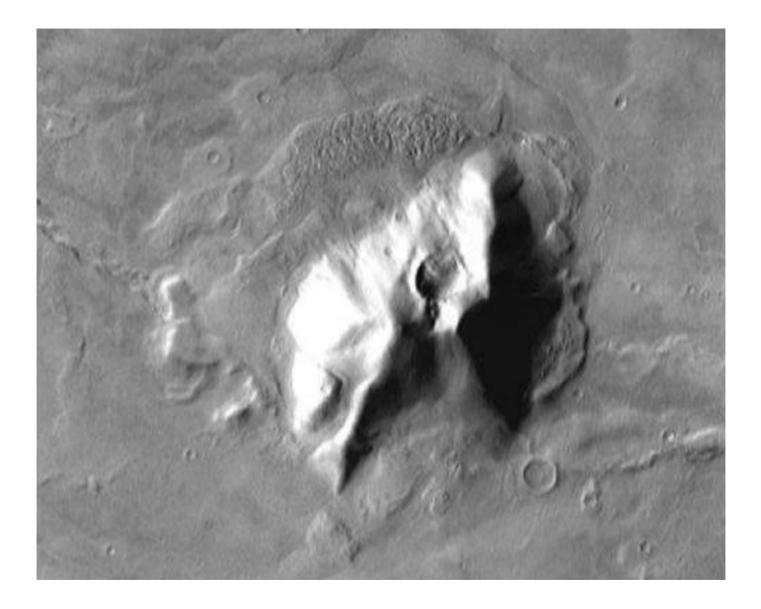


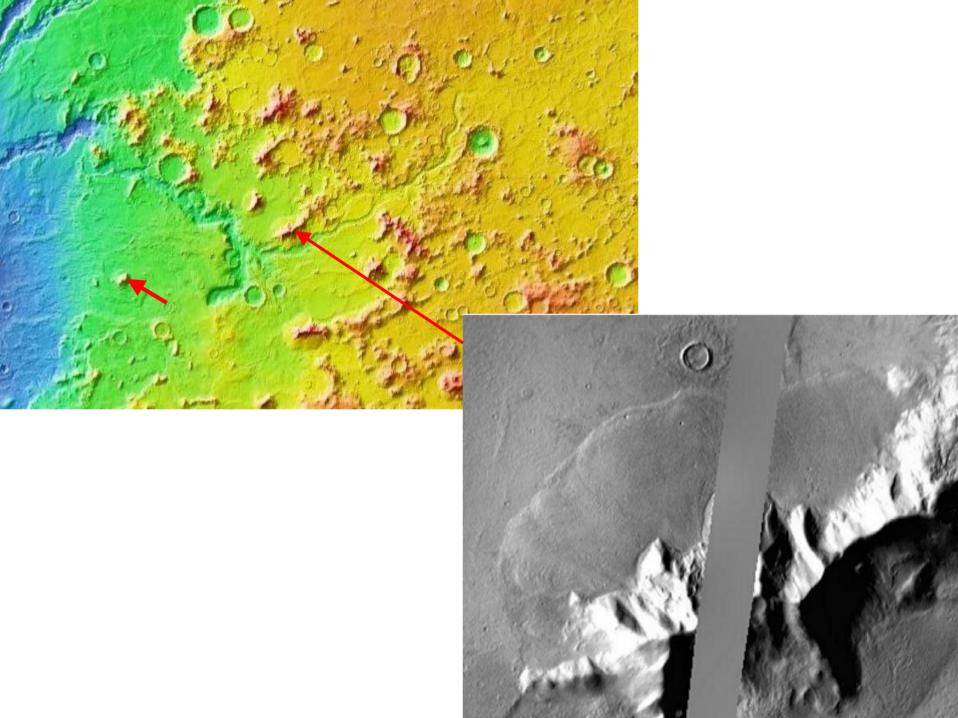
- Broad alcove represents an accumulation zone for snow and ice that incorporates debris from the massif]
- Proposed mechanism: 1) snow and ice accumulation 2) viscous flows of debris-containing ice 3) sublimation of significant volumes of ice leaving behind pitted morphology

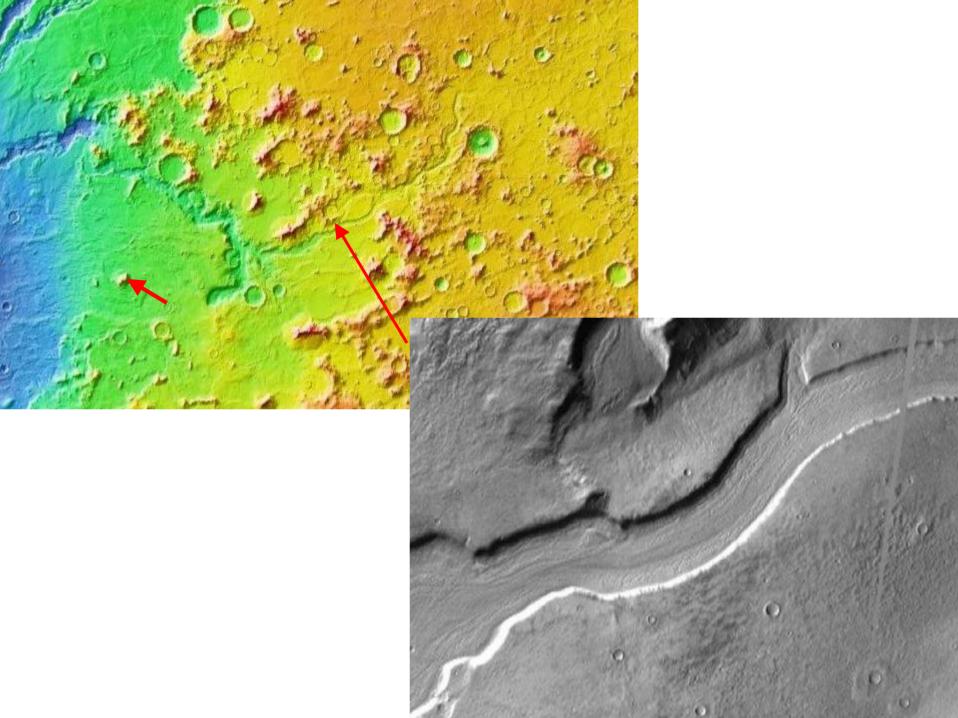


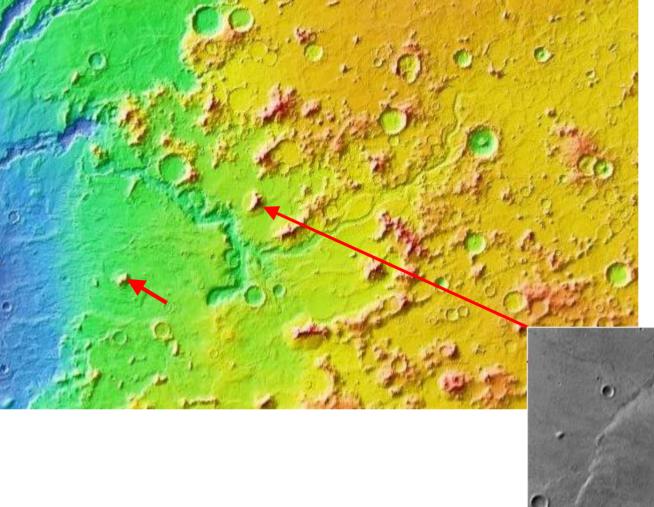


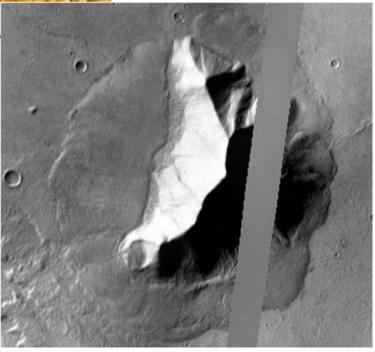


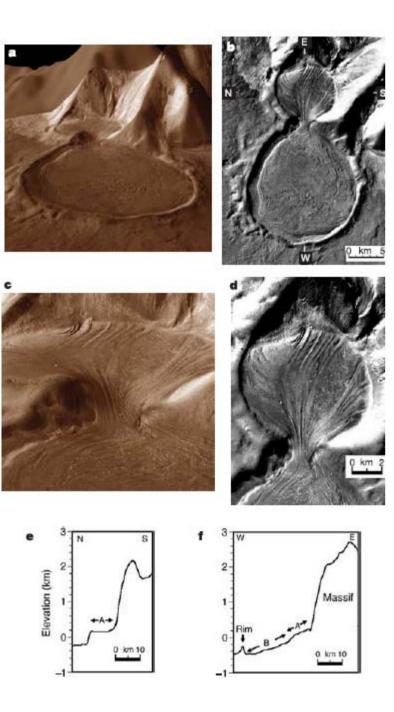










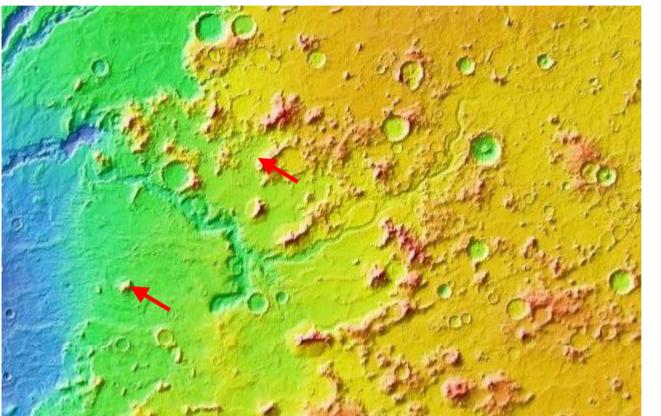


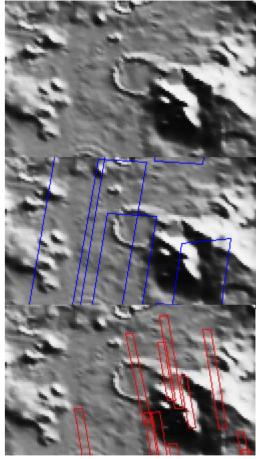
- Streamlines and lobes as evidence for viscous flow
- Streamlines start "fully formed" at the beginning of the massif
- Accumulation is done in the alcove within the massif

But, alcove is extremely steep compared with ice flow slope...

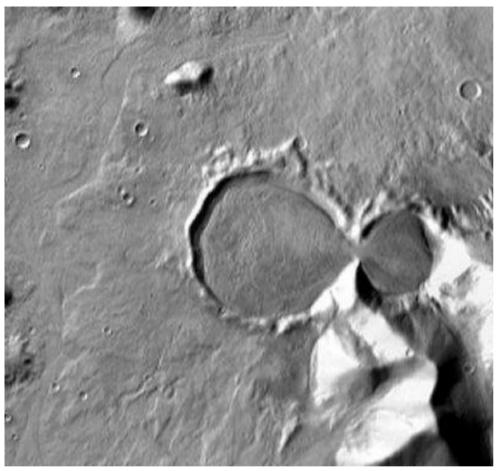
Flowlines suggest that viscosity is less than that of ice...maybe?

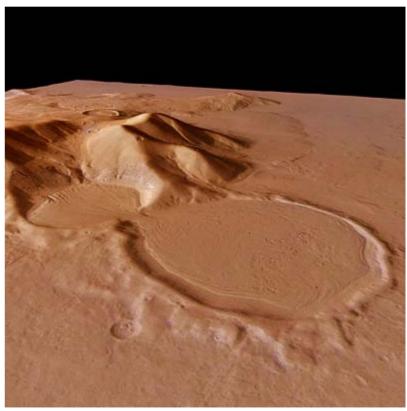
(temperatures on Mars would suggest higher viscosity ice)



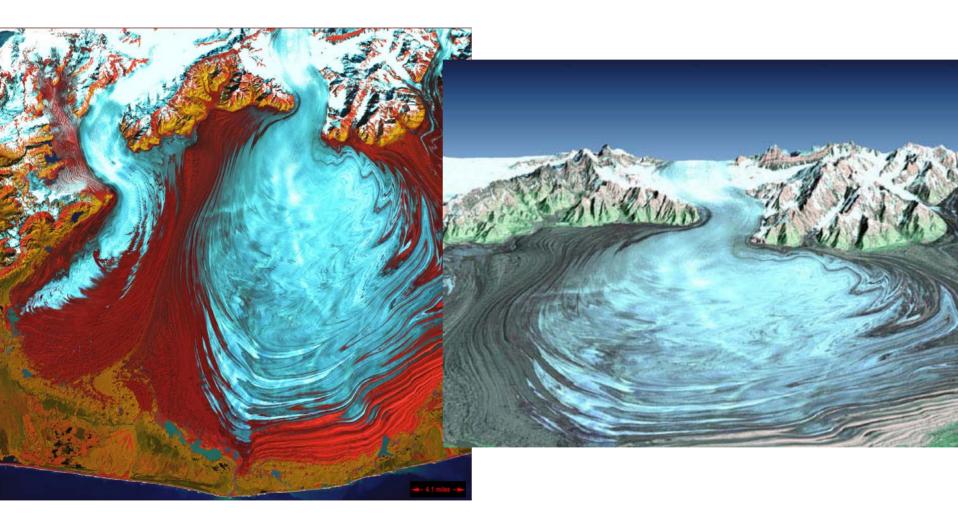


## THEMISMOC 1MOC 2

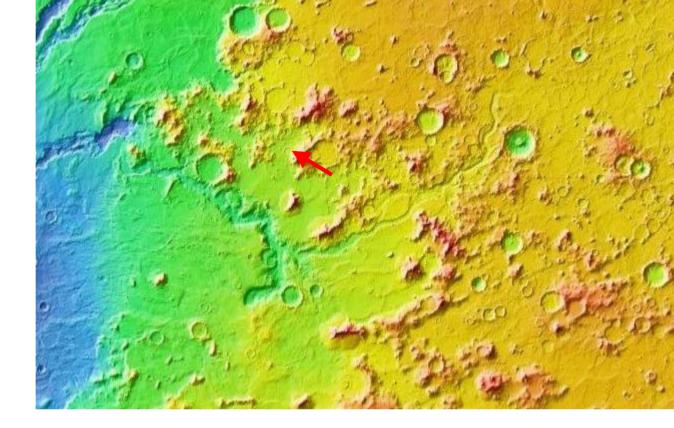


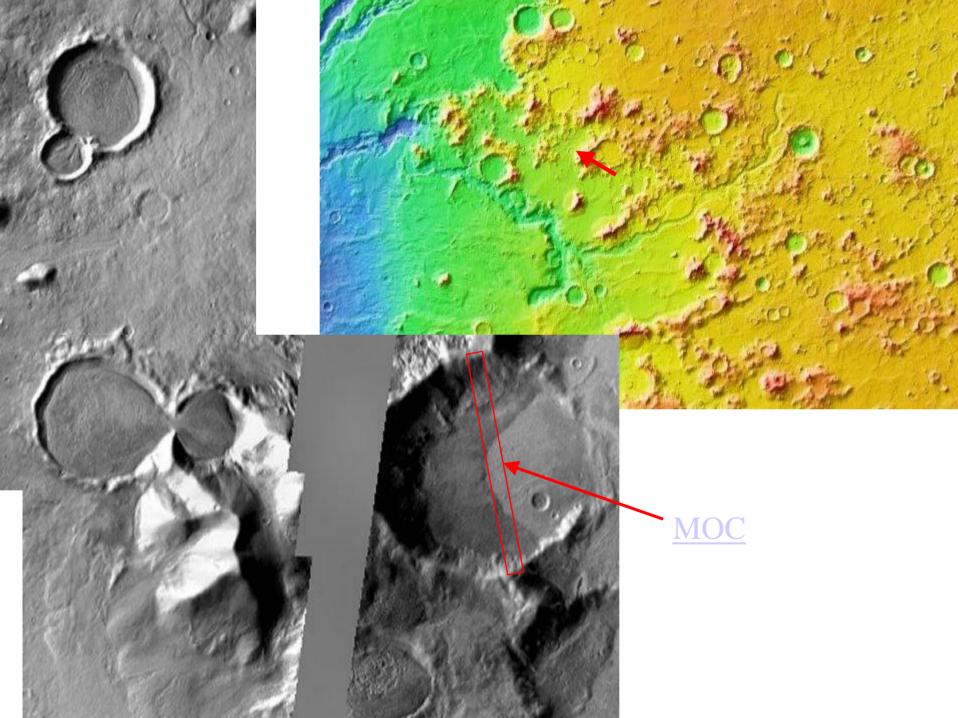


## Malaspina Glacier (Alaska): Piedmont glacier



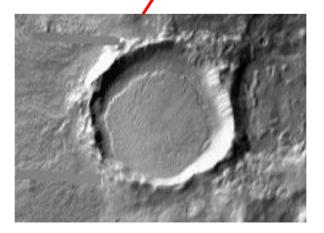




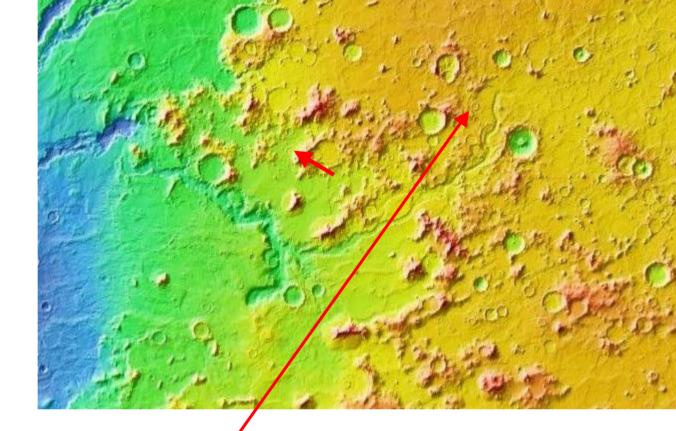










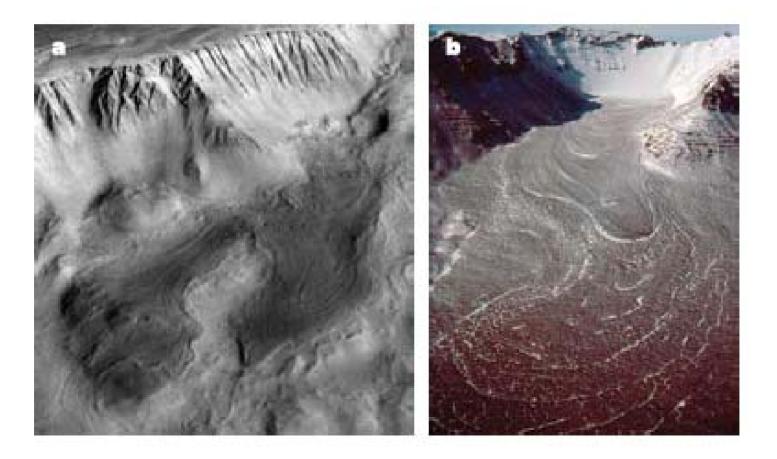


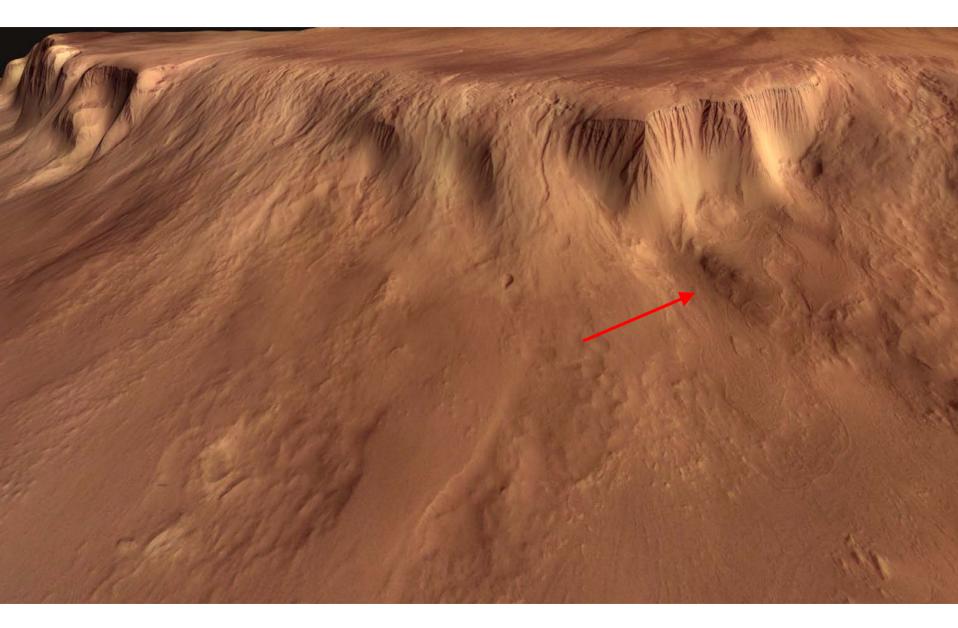


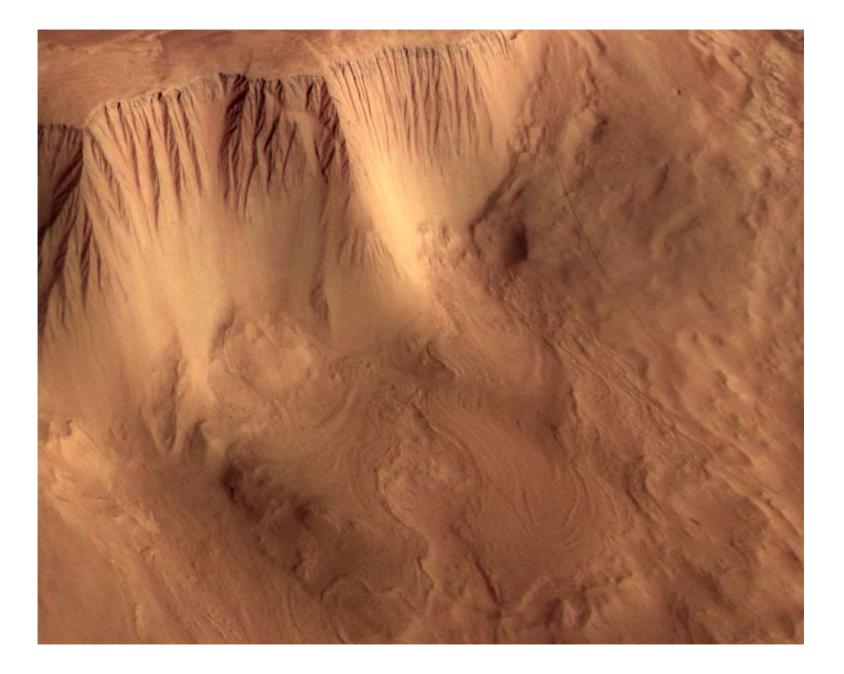
Accumulation zone?

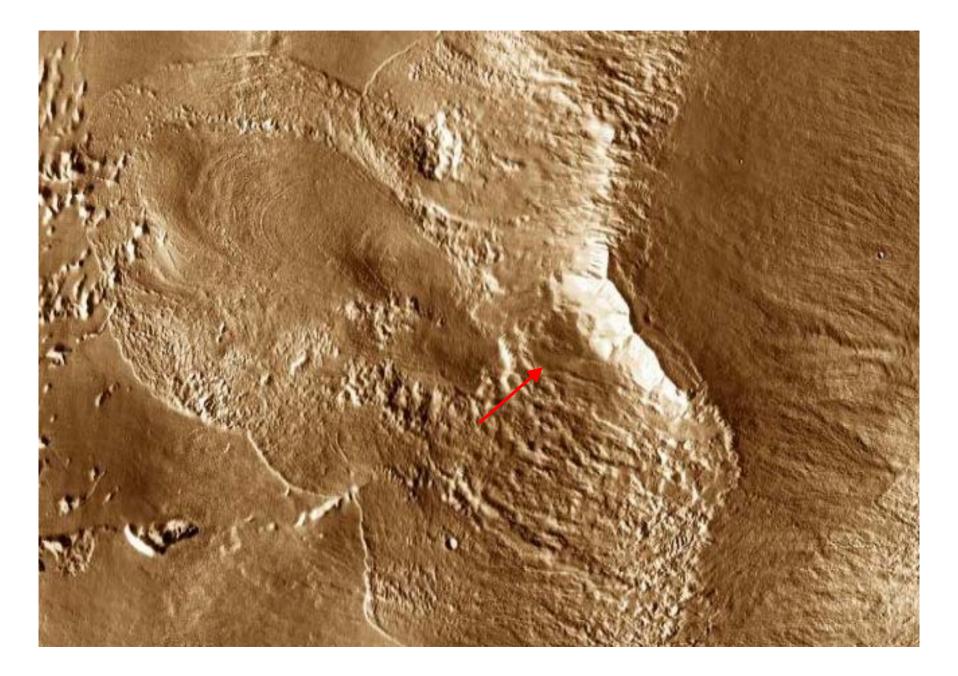
Morraines?

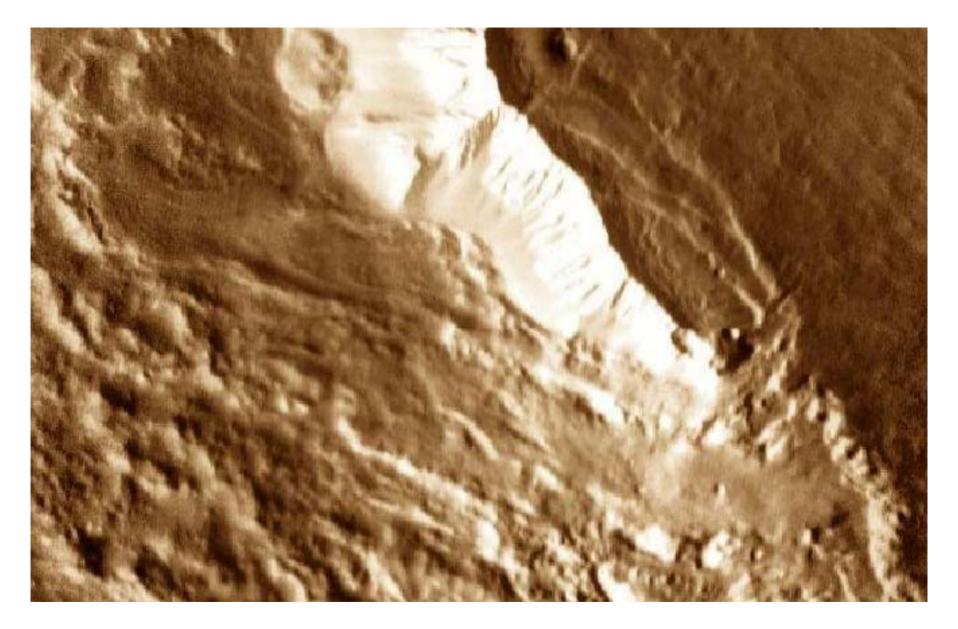
Answer: not a glacier but a mudslide!

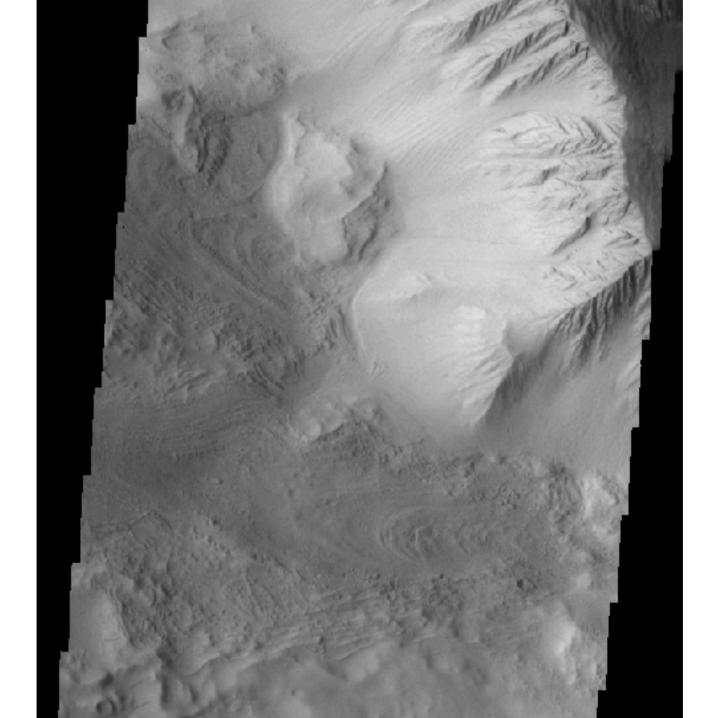












• Evidence presented by Head et al. inconclusive

•Debris Apron on Eastern Hellas massif could be ice but morphologically similar features exist on massifs without alcoves

- Viscous material within hourglass craters exist in craters that do not have associated massifs
- Glacier on Mt. Olympus is most probably a mudslide...